A toolbox for fitting non-separable space-time log-Gaussian Cox models using R-INLA

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(LEG/UFPR)







spacetime log-Gaussian Cox



One dimentional motivating example

Modeling approach

- ID code and result
- Space-time case



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Burkitt data, splancs (Rowlingson and Diggle 1993)

• Observed time lymphoma cases from year 1960 to 1975 (16 years)



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 Model: intensity function λ(t) to describe how likely a case is to happen at time t?

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Observed vs expected number of cases

sum(dat1\$year<5) # Number of observed cases (year < 5)
[1] 27</pre>

head(pred.res, 5)

year mean sd q0.025 median q0.975 smin smax cv van ## 1 0.000 0.45 0.39 0.035 0.37 1.3 0.018 1.8 0.86 0.15 ## 2 0.083 0.49 0.40 0.043 0.40 1.4 0.021 1.8 0.83 0.16 ## 3 0.167 0.53 0.42 0.059 0.45 1.5 0.025 1.8 0.80 0.18 ## 4 0.250 0.57 0.44 0.078 0.49 1.6 0.031 1.9 0.76 0.19 ## 5 0.333 0.63 0.46 0.092 0.52 1.8 0.043 2.0 0.73 0.21 sum(pred.res\$mean[pred.res\$year<5]/12) # Expected (year < 5) ## [1] 24 c(nrow(dat1), sum(pred.res\$mean)/12) # Total Obs. and Exp.

[1] 188 183

Observed data: y, n events, recorded as observed time points,
 y: t₁, t₂, ..., t_n

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Descriptive approach

• Kernel intensity estimation

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Common approach

- Aggregate the data into discrete grid cells
- Model the number of cases in each grid cell t as $Poisson(\lambda_t)$
- Model the log(λ_t)
 - pick the model framework you like the most

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Point process modeling approach

Treat the data as it is, no aggregation

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The log-Likelihood for a set of points

• Given the domain [0, L] and **y**

$$I(\lambda(.)|\mathbf{y}) = \underline{L} - \log\left(\int_0^L \overline{\lambda(t)} \partial t\right) + \sum_{i=1}^n \log(\lambda(\underline{t_i}))$$

- |L| is the size of the time domain
- $\lambda(t)$ is the intensity at time t
- t_i are time coordinates of the observed cases

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Direct likelihood approximation, Simpson et al. (2016)

$$I(\lambda(.)|\mathbf{y}) \approx L - \sum_{j=1}^{m} w_j \log \lambda(t_j) + \sum_{i=1}^{n} \log(\lambda(t_i))$$

- *m* integration points $t_1, ..., t_m$, weights $w_1, ..., w_m$
 - if the grid is equaly spaced, $w_j = w_0$

The log-Gaussian Cox point process model

• Assumption (IGCpp): log of λ follows a Gaussian process

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Figure 1: A latent Gaussian model !!!

Matern 1D field samples (smoothness ν and range)



Our model framework: Latent Gaussian models (LGM)

• Bayesian hierarchical model

$$\begin{array}{ll} \boldsymbol{y}|\boldsymbol{x},\boldsymbol{\theta} & \prod_{i} \pi(y_{i}|\eta_{i},\boldsymbol{\theta}) & \text{Several distributions!} \\ \boldsymbol{x}|\boldsymbol{\theta} & \pi(\boldsymbol{x}|\boldsymbol{\theta}) : \mathcal{N}(\boldsymbol{0},\boldsymbol{Q}(\boldsymbol{\theta})^{-1}) & \text{Gaussian (GMRF)!} \\ \boldsymbol{\theta} & \pi(\boldsymbol{\theta}) & \text{Any distribution!} \end{array}$$

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Several distributions! Gaussian (GMRF)! Any distribution!

- Precision matrix $\mathbf{Q}(\boldsymbol{ heta})$ is sparse
 - Gaussian distributions with sparse precision are Gaussian Markov random fields (GMRFs)
 - allows fast computations

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Several distributions! Gaussian (GMRF)! Any distribution!

- Precision matrix $\mathbf{Q}(\boldsymbol{ heta})$ is sparse
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 - allows fast computations
- INLA: Integrated Nested Laplace Approximations, Rue, Martino, and Chopin (2009) and Rue et al. (2017)

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The idea into practice, using inlabru

```
# d.a.t.a.
dat1 <- data.frame(year = burkitt$t/365.25)</pre>
# domain (0, 16) and integration points
mesh.t <- inla.mesh.1d(seq(-3, 20, 1/12))
# assumed GMRF model (Matern) definition
model <- year ~ 0 + Intercept +</pre>
  spde1D(map = year,
         model = inla.spde2.pcmatern(
           mesh = mesh.t.
           prior.sigma = c(1, 0.5), # P(sigma>1) = 0.5
           prior.range = c(0.1, 0.01)) \# P(range<0.1) = 0.01
# fit the model
fit.model1 <- lgcp(model, dat1)
```

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Space-time example: fox rabie cases



What do we need?

- space-time integration points
 - temporal ones and spatial ones

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- space-time integration points
 - temporal ones and spatial ones
- space-time Gaussian process
 - separable: time Kronecker space
 - non-separable: spacetime driven process

What do we need?

- space-time integration points
 - temporal ones and spatial ones
- space-time Gaussian process
 - separable: time Kronecker space
 - non-separable: spacetime driven process
- code!

Adding spatial integration points

Constrained refined Delaunay triangulation





Space-time non-separable process

• Consider the stochastic iterated heat equation, Krainski (2018)

$$\left(\gamma_t \frac{d}{dt} + L^{\alpha_s/2}\right)^{\alpha_t} u(t) = \dot{W}_{\gamma_e^2 L^{\alpha_e}}(t).$$

- differential (system of) equations with respect to time
- L: Spatial Matern
- smoothness parameters $(\alpha_t, \alpha_s, \alpha_e)$
 - flexible model, consider some cases
- scale parameters $(\gamma_t, \gamma_s, \gamma_e) > 0$

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Advantage of using SPDEs

• sparse precision matrix representation!

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Code

- consider the INLA package, Rue et al. (2017)
 - define the precision matrix using the rgeneric model
- consider the efficient PARDISO library, Kourounis, Fuchs, and Schenk (2018)

Space-time case

Preliminar result: Observed and expected cases, big areas per time



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Work in progress

- wrap it into a function for the user
 - dozens lines of code write now
- parameter interpretation \rightarrow define new ones such as
 - spatial range, temporal range, marginal variance

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References

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